

# Monitoring ice capped active Volcán Villarrica in Southern Chile by means of terrestrial photography combined with automatic weather stations and GPS

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**ABSTRACT.** Volcán Villarrica (39°25'12"S, 71°56'27"W; 2847 m a.s.l.) is an active ice-capped volcano located in the Chilean Lake District. Monitoring of the surface energy balance and glacier frontal variations, using automatic weather stations and satellite imagery, has been ongoing for several years. In recent field campaigns, surface topography was measured using Javad GPS receivers. Daily changes in snow, ice and tephra-covered area were recorded using an automatic digital camera installed on a rock outcrop. In spite of frequently damaging weather conditions, two series of consecutive images were obtained in 2006 and 2007. These photographs were georeferenced to a resampled 90 m pixel size SRTM digital elevation model and the reflectance values normalised according to several geometric and atmospheric parameters. The resulting daily maps of surface albedo are used as input to a distributed glacier melt model during a 12 day midsummer period. The spatial pattern of cumulative melt is complex and controlled by the distribution of airfall and windblow tephra, with extremely high melt rates occurring downwind of the crater and exposed ash banks. Furthermore, the camera images are also used to visualise the pattern of glacier crevassing. The results demonstrate the value of terrestrial photography to understanding the energy and mass balance of the glacier, including the generation of melt water, and the potential value of the technique in monitoring volcanic activity and potential hazards associated with ice-volcano interactions during eruptive activity.

## 29 INTRODUCTION

30 Volcán Villarrica (Figure 1, 39°25'12"S, 71°56'27"W; 2847 m a.s.l.) is considered a highly active ice-capped volcano, which is  
31 characterized in historical times mainly by mild strombolian activity (González-Ferrán, 1995; Lara, 2004), permanent degassing  
32 and periodic explosions, with the lava lake remaining at a high level (90-180 m below surface) at least since 1984 and very  
33 sensitive to the magmatic conduit activity (Calder and others, 2004; Witter and Delmelle, 2004). Concentrations of acid gases  
34 measured at the summit of the crater have been recognized as a hazard to climbers ascending the volcano, who may be exposed  
35 to concentrations above limits defined by the U.S. National Institute of Occupational Safety and Health (Witter and Delmelle,  
36 2004). Its historical eruptive activity (Petit-Breuhl and Lobato, 1994; Lara, 2004) indicates a low frequency of large explosive  
37 eruptions (Volcanic Explosivity Index, VEI between 3 and 4). More than 50 eruptive events, however, have been documented  
38 since 1558 (Petit-Breuhl and Lobato, 1994). The latest most violent eruption took place in 1971-72 when lava flows were  
39 generated, as well as 30 to 40 km h<sup>-1</sup> laharcic flows (Naranjo and Moreno, 2004) descending towards Lagos Villarrica and  
40 Calafquén (Figure 1). Lahars produced by eruptions of Volcán Villarrica in 1948-1949, 1963-1964, and 1971-1972 have resulted  
41 in the deaths of more than 75 people (Stern, 2004), and are considered the main hazard of the volcano (Moreno, 2000). The  
42 volcano is covered by a glacier of 30.3 km<sup>2</sup> (Rivera and others, 2006), mainly distributed towards the south and east where  
43 the main glacier basin (Glaciar Pichillancahue-Turbio, 17.3 km<sup>2</sup>, Figure 2), composed of partially ash/debris-covered ice is  
44 located, partially infilling a volcanic caldera depression (Clavero and Moreno, 2004). The energy balance of this glacier has  
45 been monitored since 2003 (Brock and others, 2007), and Global Positioning System (GPS) as well as Radio Echo Sounding  
46 (RES) measurements were carried out in January 2005 (Rivera and others, 2006).

47 One of the most interesting direct effects of volcanic activity on the overlying glaciers is the ice cracking or crevassing  
48 observed before or during eruptive events (Klohn, 1963; Fuenzalida, 1976; González-Ferrán, 1995; Fuentealba and others,  
49 1985). Ice cracking has been detected by seismicity (Metaxian and others, 2003), however here we propose to observe possible  
50 ice cracking by obtaining daily photographs of the ice. This process could be important for ice flow and for the hydrological  
51 balance of the glacier, as crevasses are the main pathway for meltwater to enter the glacier's en- and subglacial drainage system  
52 (Fountain and others, 2005; Fountain and Walder, 1998).

53 A second direct effect of volcanic activity is the deposition of pyroclastic materials on top of the glaciers, changing the albedo  
54 and affecting the energy balance by providing insulation from atmospheric heat and insolation where particles are large or  
55 the cover is continuous (Adhikary and others, 2002; Kirkbride and Dugmore, 2003). Studies of tephra thermal properties and  
56 glacier melt rates at Volcán Villarrica (Brock and others, 2007), identified a very low thermal conductivity of 0.35 W m<sup>-1</sup> K<sup>-1</sup>  
57 in the lapilli tephra which blankets most of the ablation zone. Furthermore, a critical thickness of just 5 mm of tephra cover was  
58 found to reduce the melt rate of the buried ice compared with a bare surface. Consequently, volcanically produced materials  
59 appear to have a large positive impact on the mass balance of glaciers on Volcán Villarrica, due to an extensive mantle of  
60 insulating tephra in the lower ablation zones. In this study quantitative assessment of the total area of tephra cover and  
61 tephra-free ice slopes and snow albedo have been made through terrestrial photography.

## 62 AIMS AND METHODS

63 The main aim of this paper is to analyse albedo variation and its impact on melt on Glaciar Pichillancahue-Turbio of Volcán  
64 Villarrica by means of terrestrial photography. Previous measurements of the glacier areal variations (Rivera and others, 2006)  
65 have been updated until year 2007 which permitted the assessment of recent ice retreat on the volcano's glaciers.

## 66 Meteorological data

67 The meteorological data were obtained by an automatic weather station (AWS) that was located during the summer on the  
68 surface of the glacier at 1933 m asl, near the Equilibrium Line Altitude, located at  $\sim 2000$  m a.s.l. during the year 2003/4  
69 (Rivera and others, 2006). During the winter the AWS was moved to a location on a rock outcrop at 1925 m a.s.l., next to  
70 the fixed camera (Figures 2 and 3). The AWS recorded incoming and reflected shortwave radiation, net all-wave radiation, air  
71 temperature, air humidity and wind speed at 2 m above the surface with hourly mean values recorded on a Campbell CR10  
72 data logger. Albedo was measured using a Kipp and Zonen CM6B albedometer sensitive to radiation in the range 0.3 - 2.8  
73  $\mu\text{m}$ , with the sensors mounted in a surface parallel plane. Details of the collected variables and technical details of the utilized  
74 sensors at the AWS are described in Brock and others (2007).

## 75 Oblique photography

76 For measuring albedo, surface changes and tephra cover, an automatic camera (Canon EOS 300D) was installed at the upper  
77 part of the rim surrounding the main volcano edifice (upper part of the caldera, Figure 2) from where daily photographs of the  
78 glacier were obtained. The camera has a 6.3 Megapixel CMOS sensor and recorded the images in a 2 GB flash card memory.  
79 It was fitted with a high quality, fixed 24 mm focal length lens to minimise optical distortions. This camera sensor has some  
80 sensitivity in the near infrared spectrum, at least to 1000 nm, and beyond that point if the IR filter is removed (experimental  
81 tests by the authors). The camera was inserted into a Pelikan sealed box where it was controlled by a Canon timer. The system  
82 was powered by 12-V batteries which were fed by a solar panel installed nearby.

83 Conventional photography is a powerful medium for collecting and storing information. If this information can be located  
84 precisely in space, then photography becomes a powerful tool for quantitative analysis. Here we use a tool for georeferencing  
85 oblique photography developed by Corripio (2004), using a single image and a digital elevation model (DEM). The accuracy  
86 of the technique will depend on the accuracy of the DEM and on the quality of the photographic image, especially the degree  
87 of distortion and aberration produced by the lens. This technique does not produce elevation data. It actually requires an  
88 existing DEM. What the tool does is to locate the geographical position of every pixel in a photographic image. It is therefore  
89 useful to map land cover and to assess surface cover change. This technique follows standard procedures for perspective  
90 views in computer graphics or photogrammetry (Fiume, 1989; Foley and others, 1990; Slama and others, 1980). Basically it  
91 consists in creating a virtual photography of the DEM that can then be scaled to the resolution of the photographic image to  
92 establish a mapping function between pixels in the photograph and grid cell points. This allows locating the exact position of  
93 pixels in the oblique image. The georeferencing process consists of a viewing transformation applied to the DEM in which the  
94 coordinates of every grid cell are firstly translated to refer them to a coordinate system with origin at the camera position.  
95 Then a transformation is applied according to the viewing direction and focal length of the camera. This results in a three  
96 dimensional set of points corresponding to the cells in the DEM as seen from the point of view of the camera. Finally, the  
97 resulting viewing transformation is projected into a two-dimensional viewing 'window', corresponding to the area of the film  
98 and scaled proportionally. The process is explained in detail by Corripio (2004), it has been coded in IDL under the Creative  
99 Commons license and it is freely available from the authors.

100 Using this technique the evolution of the snow cover was mapped accurately during the acquisition periods. Once the  
101 photograph is georeferenced, the reflectance values are normalized according to the viewing geometry, the angle of incidence  
102 of sun's rays on the slope, the ratio of direct to diffuse radiation, the atmospheric transmittance between the pixel location  
103 and the position of the camera, and the effect of radiation reflected from the surrounding slopes. The final result is a map of  
104 normalized reflectance values, or relative albedo. The atmospheric transmittance was calculated using the radiative transfer

105 model MODTRAN (Berk and others, 1989) and general knowledge of the local atmospheric conditions from the AWS. The  
106 photographic image was mapped to a resampled DEM from the Shuttle Radar Topography Mission (SRTM, acquired by  
107 JPL/NASA in 2000). It is important to notice that the resampling procedure to a 10m resolution DEM cannot increase the  
108 actual resolution of the original DEM, at 90m, but allows the extraction of more information from the photography within the  
109 known spacial limits of the original DEM grid cell. The DEM was not contemporary to the photographic image acquisition.  
110 This may add some errors to the results, as the slope and aspect may be slightly different from one date to another, and the  
111 snow accumulation may change. It is completely out of our financial possibilities to acquire a high resolution DEM for every  
112 campaign, but we are working on the design of some terrain measurements that allow us a precise evaluation of the errors  
113 incurred by using slightly old digital elevation models, such as simultaneous measurements of albedo on different points and  
114 comparison with those derived from georeferenced photography.

115 Unfortunately the instrumentation for this project was a constant battle against the elements. Despite testing the kit under  
116 laboratory conditions, the camera case was twice flooded by severe weather conditions, which also prevented collection of data  
117 on one occasion. The camera setup was finally destroyed by a lightning strike in March 2007.

### 118 **GPS survey**

119 Several kinematic and static GPS surveys were conducted to map the extent of the glacier and to georeference tie points for  
120 photogrammetric purposes (Figure: 3). Javad's dual-frequency GPS receivers and antennas were used exclusively applying a  
121 sampling rate of at least 2 seconds and an elevation cut-off angle of 10 degrees. At all times a nearby reference station mounted  
122 on bedrock was occupied to provide geodetic quality measurements for differential positioning at a centimeter level. Thereby,  
123 baseline lengths never exceeded a few kilometers. GPS data were postprocessed using Waypoints GrafNav Softwarepackage  
124 version 7.70 where precise ephemeris and clock information provided by the International GNSS Service (IGS, final products)  
125 were incorporated. In January 2005 the local reference station was linked to the SIRGAS (Sistema de Referencia Geocéntrica  
126 para las Américas) which is the regional realization of the International Terrestrial Reference Frame.

### 127 **Satellite imagery**

128 In order to update the glacier variations at the volcano, several satellite images and aerial photographs were used (Table 1).  
129 All the satellite images were geo-located and orthorectified using the regular IGM cartography and available Digital Elevation  
130 Models (e. g. SRTM and AIRSAR), (Rivera and others, 2006). Once the satellite images were orthorectified, classification  
131 procedures based upon spectral band ratios were applied to account for the glacier extent and snow/ice/debris classification  
132 (Paul and others, 2002). Aerial photographs were stereoscopically analyzed and the resulting information was transferred with  
133 a Zoom Transfer Scope (ZTS) to the regular cartography, with ice fronts being digitally compared to the satellite (Benson  
134 and Follet, 1986). All glacier limits were analyzed using GIS commercial software, such as IDRISI 32 for Windows, Arc-Info  
135 version 8.0.1 and PCI Geomatica, allowing an accurate estimation of areas and frontal changes (Figure 2).

### 136 **Energy balance and melt modelling**

137 We run a modelling experiment to assess the effect of tephra deposition on the glaciated surface of the volcano and its  
138 influence on mass balance and runoff. Using meteorological data for twelve days in January, 2007 and the corresponding  
139 albedo variations derived from the photographic images, a distributed energy balance model was applied to eastern slopes of  
140 the Volcán Villarrica. Thus, we could assess the effect of diminished albedo due to volcanic ash deposition on the mass balance  
141 and melt-water production from the volcano glacier. In January 2007 there were no available wind speed data from the AWS.  
142 These data were derived from the NOAA archived 1 degree resolution AVN model outputs for the grid cell corresponding to

143 Volcán Villarrica. The data were chosen at a 750 hPa pressure level, which corresponds to the middle height of the volcano's  
144 cone. The model applied (SnowDEM-Snow Distributed Energy balance Model) is explained in detail in Corripio (2003a). It  
145 is a distributed, multilayered energy balance model, which takes into consideration radiative fluxes, heat interchange with the  
146 atmosphere, evaporation or sublimation and heat flux due to precipitation. Short wave is evaluated according to a detailed  
147 parametric radiative transfer model of the atmosphere plus terrain effects (Corripio, 2003b, 2004; Strasser and others, 2004).  
148 Longwave is estimated from atmospheric humidity and temperature plus terrain parameters and ground/snow emissivity  
149 according to geological characteristics if data are available. Tests on complicated surfaces in the Andes such as penitentes,  
150 testify the ability of the model to reproduce snow surface temperature (Corripio and Purves, 2005). Application to a watershed  
151 in the Alps for the estimation of meltwater runoff gave values within 6% of measured runoff for a runoff gauge with a 10%  
152 accuracy (unpublished data).

## 153 RESULTS AND DISCUSSION

154 The frontal and areal glacier variations of Glaciar Pichillancahue-Turbio have been updated until year 2007 (Table 2). Both  
155 main arms of the glacier have continued retreating at similar rates to frontal length changes measured in recent decades  
156 (Casassa and others, 2004). These glacier tongues are totally debris covered and only at steep flanks is bare ice visible due  
157 to the backwasting ablation process (figure 2). In spite of the insulation provided by the ash and debris covering the ice, the  
158 glacier has lost an important area in recent decades, much higher than other debris free glaciers also located on top of active  
159 volcanos (Rivera and others, 2006). The present extent of Villarrica's glaciers are shown in figure 2, while recent variations  
160 are summarised in table 2.

161 In order to monitor the glacier at a daily resolution, we used oblique terrestrial photographs that were georeferenced to a  
162 DEM with the help of accurate ground control points (GCP) measured on the glacier surface, as shown in Figure 4. This tool  
163 can be applied to precisely locate snow features on the surface of the glaciers in areas that are of very difficult access. The  
164 eastern upper side of the volcano is constantly swept by ash falls and toxic gases from the crater fumaroles. This made direct  
165 surveying impossible without appropriate protective clothes and breathing masks. Thus remote sensing is a more convenient  
166 alternative as shown in Figure 5. This figure shows the details of two different images and the changing position of crevasses on  
167 the upper section of the volcano. The left image is 25 March 2006 and right image is 14 January 2007. The derivation of flow  
168 rates and surface variations using this approach is currently under study and results will be given in a future publication. This  
169 preliminary demonstration is shown to illustrate the potential of this technique for high temporal resolution surface monitoring  
170 in hazardous or difficult to access environments.

171 The pattern of tephra dispersion is clearly visible on the georeferenced image of the volcano on 25/12/2005 (Figure 6). The  
172 conical shape of the georeferenced image is due to the field of view of the camera. There is a thick tephra layer around the  
173 crater rim and a band of darker snow to the SE following the prevailing W and NW winds. The amount and area of tephra  
174 deposition depends on the intensity of fumarolic activity and the concurrent winds. It would be possible to model windflow  
175 across the volcano, however it is far more difficult to predict the spatial and temporal variability of fumarolic activity. It is  
176 therefore very difficult to anticipate the ash dispersion pattern. The only solution is to observe it at relatively high temporal  
177 resolution and incorporate the results to any model of the glacier surface.

178 Here we present the results for a set of eleven images during a clear sky period from January fifth to January sixteenth,  
179 2007. The albedo of the visible area of the volcano was derived from the photographic images and incorporated into the  
180 energy balance model. The albedo derived from the images reveals increasing values toward the fringes of the visible area (the

181 northern and southern slopes). We believe this is a realistic result, as those slopes are away from the prevailing winds and  
182 suffer less tephra pollution by volcanic fallout. In fact, a visual inspection of the northern slopes while skiing Villarrica a week  
183 earlier showed thick ash layers on the uppermost section of the volcano near the crater and clean, metamorphosed granular  
184 snow below that point to an altitude of about 1800 m a.s.l. Snow was fine grained on the upper section, interspersed with small  
185 wind deposits of highly broken precipitation particles (Colbeck and others, 1990), on the lower section it consisted of larger  
186 snow grains with high water content.

187 Previous work suggests that the critical ablation threshold is passed as soon as the tephra forms a continuous layer at the  
188 surface, with insulation and melt reduction overriding the influence of albedo lowering (Brock and others, 2007). In this analysis  
189 we assumed the critical debris thickness to correspond with an albedo of 11% (based on the broadband albedo of andesitic  
190 basaltic tephra, ASTER spectral Library, <http://speclib.jpl.nasa.gov/>), below which snow melt is reduced relative to bare  
191 snow surfaces. Spurious high values may be due to errors in the precise boundary between tephra covered and bare snow areas.  
192 These errors are of the order of two pixels of the original DEM resolution at the borders of the image, or  $\pm 180$ m. Areas  
193 where the viewing angle is very shallow have been masked, but sub-pixel variation in slope may add to this error. Increasing  
194 the precision without the use of photogrammetric cameras and a very up to date DEM is unlikely. The resulting ablation map  
195 (Figure 7) shows high spatial variability. This variability may be enhanced by post-depositional reworking of the tephra layers,  
196 and facilitated by a positive feedback, as surface particles will tend to aggregate while melting on concave surfaces (Drewry,  
197 1972), which are initially caused by differential melt. Of particular note are areas of locally enhanced melting downwind of  
198 exposed ash banks in the lower third of the image, and accelerated melting over large areas downwind of the crater, which are  
199 exposed to almost continuous airfall tephra deposition. These irregularities are also discernible on any transect along and across  
200 the eastern slopes of the volcano, as shown in figure 10. This figure shows the modelled differences in accumulated ablation for  
201 a transect along the 2500 m elevation contour and along a transect from the summit crater towards the camera standpoint.  
202 Along the elevation isopleth the pattern is approximately symmetrical, with minimum values towards the slopes that are less  
203 subject to ash deposition and a maximum on the SE slopes. The horizontal gradient is rather large, with differences in ablation  
204 of several cm over a few tens of meters. On the vertical transect, we can observe a clear anomaly, as ablation increases with  
205 elevation and reaches a local maximum near the crater (distance 0 to 500 m), where ash depositions are more intense. It then  
206 follows an irregular but decreasing trend downslope, as ashes get more dispersed with distance from the source. There are two  
207 peaks about 5 km from the crater associated with older ashes that are resurfacing after the overlying snow has disappeared.  
208 In these areas tephra is probably thick enough to insulate the underlying ice and reduce melting. Brock et al. (2007) measured  
209 a mean daily melt rate of 46 mm w.e. at 3 stakes set on snow with mean albedo 0.51 in the vicinity of the AWS, under  
210 similar meteorological conditions in the second half of January 2004. The measured melt rate (12 day cumulative melt = 506  
211 mm) corresponds well with modelled values using photographic derived albedo in January 2007 over large areas of snow with  
212 relatively light tephra cover (Figure 7 and Figure 10).

213 The effects of variable deposition, and redistribution of fallen tephra, on spatial patterns of melt are illustrated by comparison  
214 with modelled melt rates when these spatial variations are neglected. Figure 8 shows the differences in modelled cumulative  
215 melt rates for the same period, replacing photographic derived albedo with the Brock and others (2000) ageing-curve albedo  
216 parameterisation. This parameterisation calculates albedo as a function of accumulated temperatures since the most recent  
217 snowfall and assumes albedo decay is caused by snow metamorphism and the build up of impurities over time, but does not  
218 account for spatial variation in snow impurity content. While the model with parameterised albedo is able to account for some  
219 of the along glacier (vertical) variation in melt rates associated with slower snow metamorphism at higher elevations, cross

220 glacier variation in melt rates, other than that due to aspect and shading is not incorporated and, in particular, the high  
221 spatial variability in melt rates immediately below the crater and in the vicinity of exposed ash banks is missed (Figure 8).  
222 This demonstrates that using photographic derived albedo provides an improved depiction of the actual spatial variation of  
223 surface albedo.

224 The differences in cumulative ablation between models accounting or not for tephra deposition reach peak values over 40  
225 cm in some spots, however these peak values are probably an overestimation. The errors in this computation are due to a)  
226 subpixel slope and albedo variation; b) local areas of very shallow viewing angle where the average DEM slope is steeper and  
227 c) projection of cones and mounts into adjacent grids due to almost parallel viewing angle. Therefore only the areas where  
228 the viewing angle of camera is over 60 degrees are reliable. This results in differences in ablation of the order of 20 to 30 cm  
229 between model runs considering ash cover or clean snow. While at the end of the accumulation season the surface of the glacier  
230 is relatively smooth, at the end of the ablation season it presents many concavities a few meters depth and a few tens of meters  
231 in diameter (Figure 9). The negative values are areas where the photography derived albedo is higher than the parameterised  
232 albedo. This is likely to happen in small concave areas which get more shading from the sun. These small concavities are  
233 not registered in the DEM, and therefore a more precise estimation of their albedo would require a higher resolution DEM.  
234 The irregular snow surface is likely to be due to a combination of effects by glacier flow and crevasse formation together with  
235 differential ablation.

236 The increase in ablation of snow due to tephra impurities was estimated by Brock and others (2007) to be 8% at the  
237 location of the AWS. The same authors recorded rates of melting similar to those produced here for the areas with higher  
238 ash depositions. Ablation seems certainly higher further up the slope, as the albedo is lower. This can also be seen from the  
239 longitudinal transect (Figure 10). The average modelled increase in ablation for the pixels visible from the camera viewpoint is  
240 13.6% in this study. It is questionable whether this increased ablation and related meltwater production, might have an effect  
241 on the glacier dynamics, given the high porosity of the ground. It is, however, likely that it has an impact on the ground water  
242 recharge and on runoff further down the valley, with additional implications for the ecological system (e.g. Hauer and others,  
243 1997). A network of piezometers together with runoff gauges would be the ideal instrumentation to detect this effect on water  
244 levels, and hopefully will be incorporated in future research programs in the region.

## 245 CONCLUSIONS

246 A number of photographs were obtained from a fixed camera installed at Volcán Villarrica. These photographs have been used  
247 to incorporate daily spatial albedo variations into a glacier melt model based on AWS data, SRTM elevation data and GPS  
248 measurements on the ground. The modelled ablation shows significantly higher values than previously measured on the volcano  
249 (Brock and others, 2007) in areas directly downwind from ash sources, i.e. the crater and exposed tephra banks. These ashes  
250 reduce albedo and as a result increase the surface melt rate. The maximum values were found near ash/debris hummocks where  
251 winds are blowing material onto the snow. However, large areas of the glacier experience high ablation (low albedo) in the area  
252 located downstream to the E and SE from the crater, where airfall tephra are frequently deposited. Another interesting result  
253 obtained here, is the possibility of detection of crevassing at the flanks of the volcanic cone. Changing patterns of crevasses  
254 could indicate subglacial volcanic activity, as well as glacial ice dynamics. Further research will be needed to extract surface  
255 velocities using, e.g. feature tracking in the daily photographs. The frontal variations of the glacier have also been updated,  
256 confirming previous receding trends, in spite of the thick insulating debris which mantles most of the glaciers' ablation zones.  
257 Continued glacier retreat in spite of the insulating tephra is likely to be due to a combination of reduced accumulation due

258 to decreasing precipitation trends in the region (Bown and Rivera, 2007), rapid melt at exposed ice cliffs along crevasses and  
259 possibly enhanced basal melting in areas of high geothermal heating (Rivera et al, 2006).The role of ash deposition on top of  
260 the snow surfaces has been discussed here, but more detailed analysis of the mass balance of the glacier needs to be done in  
261 order to understand the specific driving factors explaining the retreating trend, especially how the decrease in precipitation  
262 observed in the Chilean Lake District (Bown and Rivera, 2007) is affecting the accumulation on the glacier.

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**Table 1.** Aerial photographs and satellite images

	OEA*	LANDSAT MSS**	SAF*	SAF*	GEOTEC*	ASTER**	ASTER**	ASTER**
DD/MM/YY	13/12/1961	08/02/1976	07/01/1983	15/12/1987	30/01/1988	13/02/2003	02/02/2005	24/02/2007
Scale / Resolution (m)	1:50 000	57	1:30,000	1:50,000	1:70,000	15	15	15

\* Aerial photographs

\*\* Satellite image

**Table 2.** Frontal and areal changes of Glaciar Pichillancahue-Turbio

	Period	Length change (m)	Annual length change ( $\text{m a}^{-1}$ )	area change ( $\text{km}^2$ )	Source
<b>Glaciar Turbio</b> <b>-39°25', -71°53'</b>	1961-1983	-407	-19	-0.02	Casassa <i>et al</i> , 2004
	1983-1987	-43	-9	-0.08	
	1987-1998	-200	-20	-0.35	
	1998-2003	-679	-135	-0.43	This study
	2003-2005	stable		-0.14	
	2005-2007	-60	-30	-0.05	
		<b>1961-2007</b>	<b>-1389</b>	<b>-30</b>	<b>-1.07</b>
<b>Glaciar Pichillancahue</b> <b>-39°26', -71°52'</b>	1961-1983	-288	-14	-0.93	Casassa <i>et al</i> , 2004
	1983-1987	-51	-11	-0.11	
	1987-1998	-414	-41	-1.05	
	1998-2003	-52	-10	-0.51	This study
	2003-2005	-136	-68	-0.24	
	2005-2007	-95	-48	-0.20	
		<b>1961-2007</b>	<b>-1036</b>	<b>-23</b>	<b>-3.03</b>

271 FIGURES

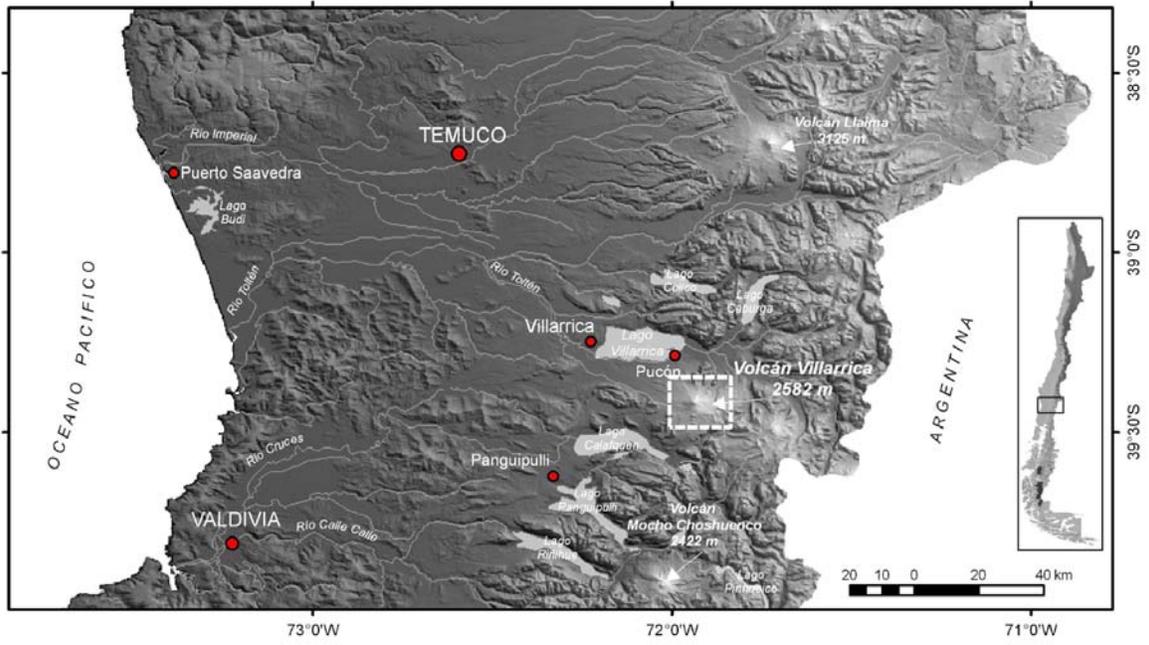
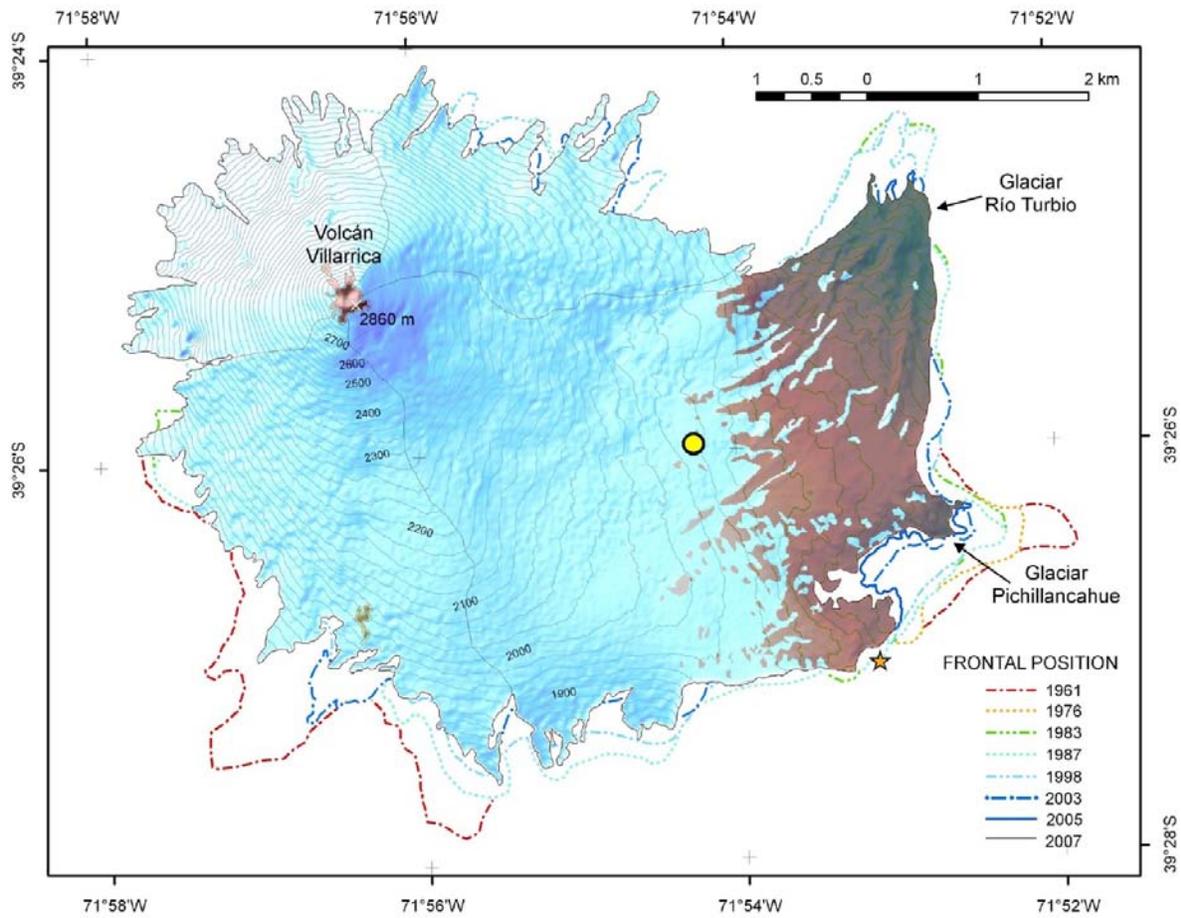
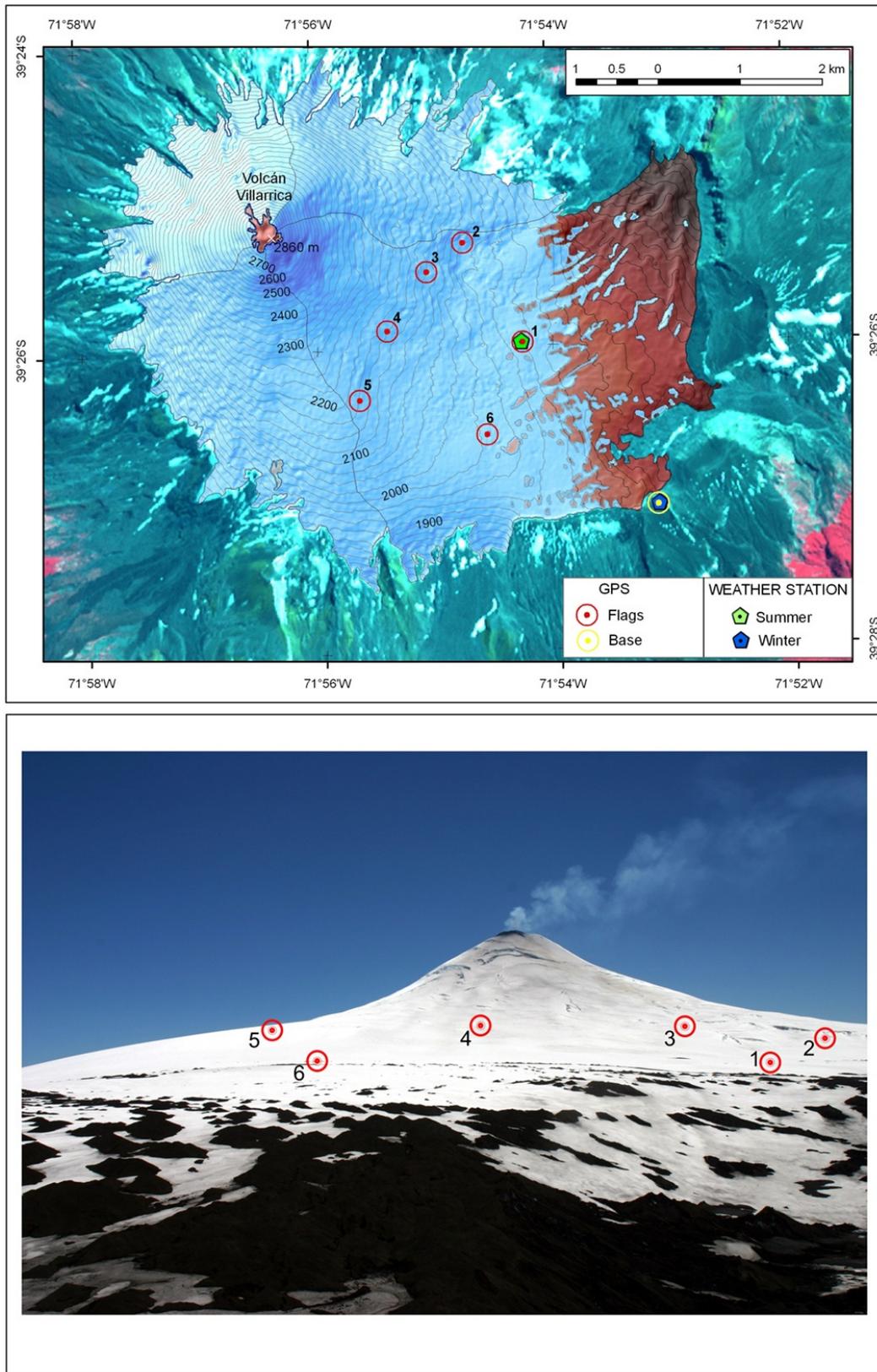


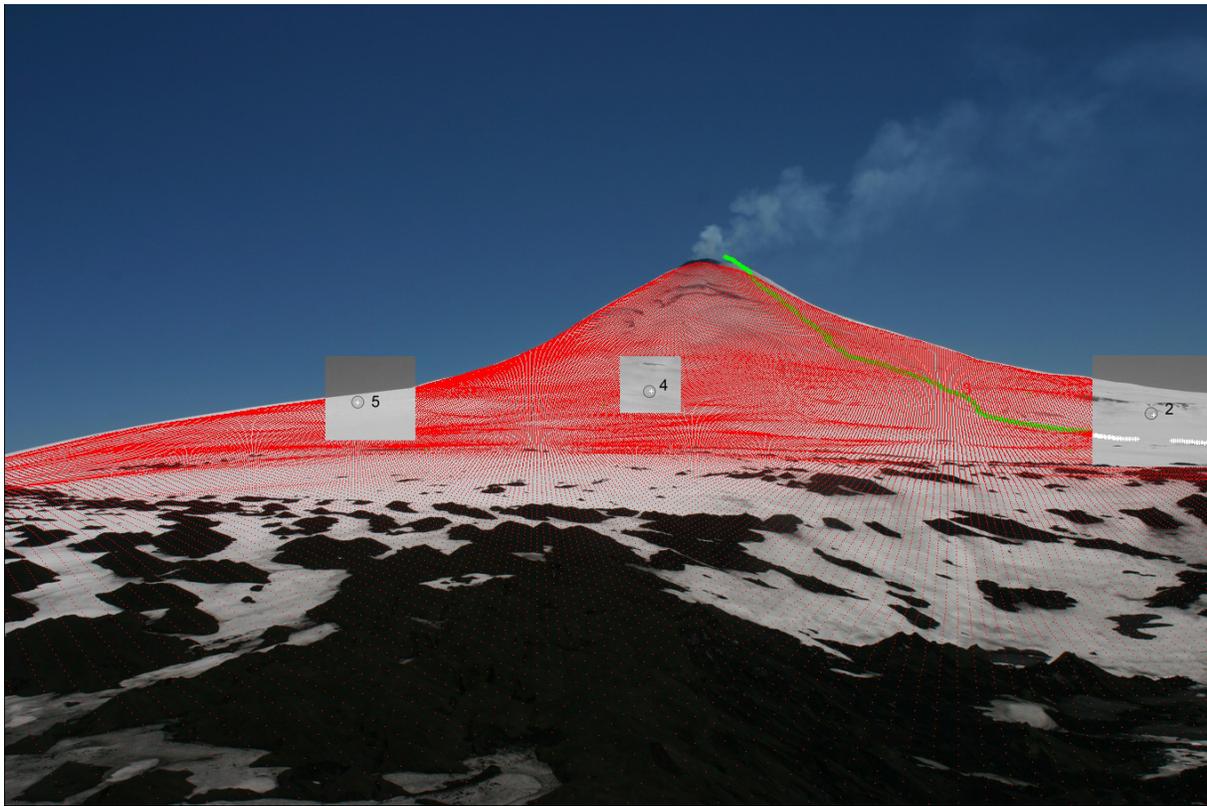
Fig. 1. General location map of Volcán Villarrica



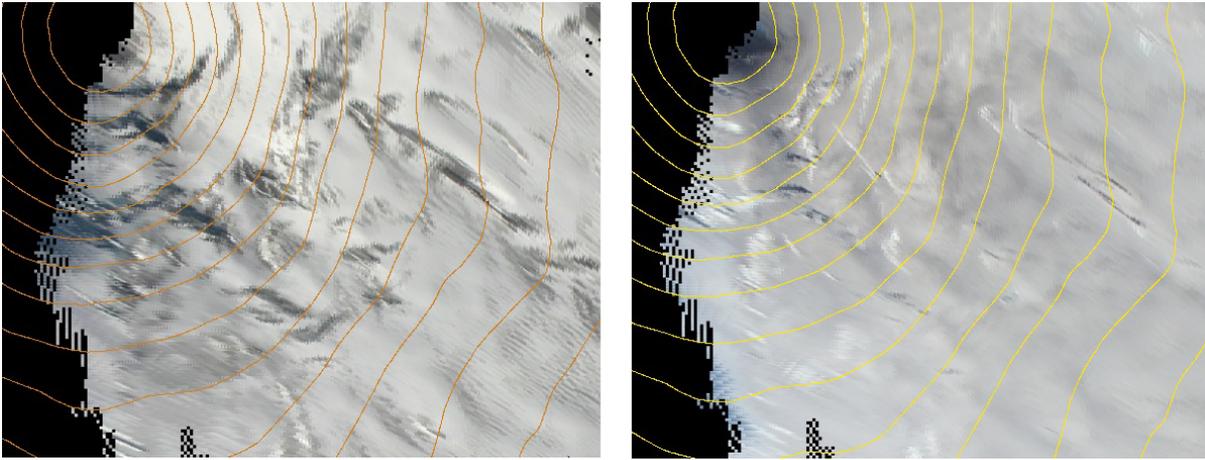
**Fig. 2.** Glacier variations 1961-2007 at Volcán Villarrica. The yellow dot shows the summer AWS location. The star is showing the location of the Camera, the GPS base Station and the AWS in winter. Blue is snow or ice cover, while brown shows areas of permanently tephra-covered ice.



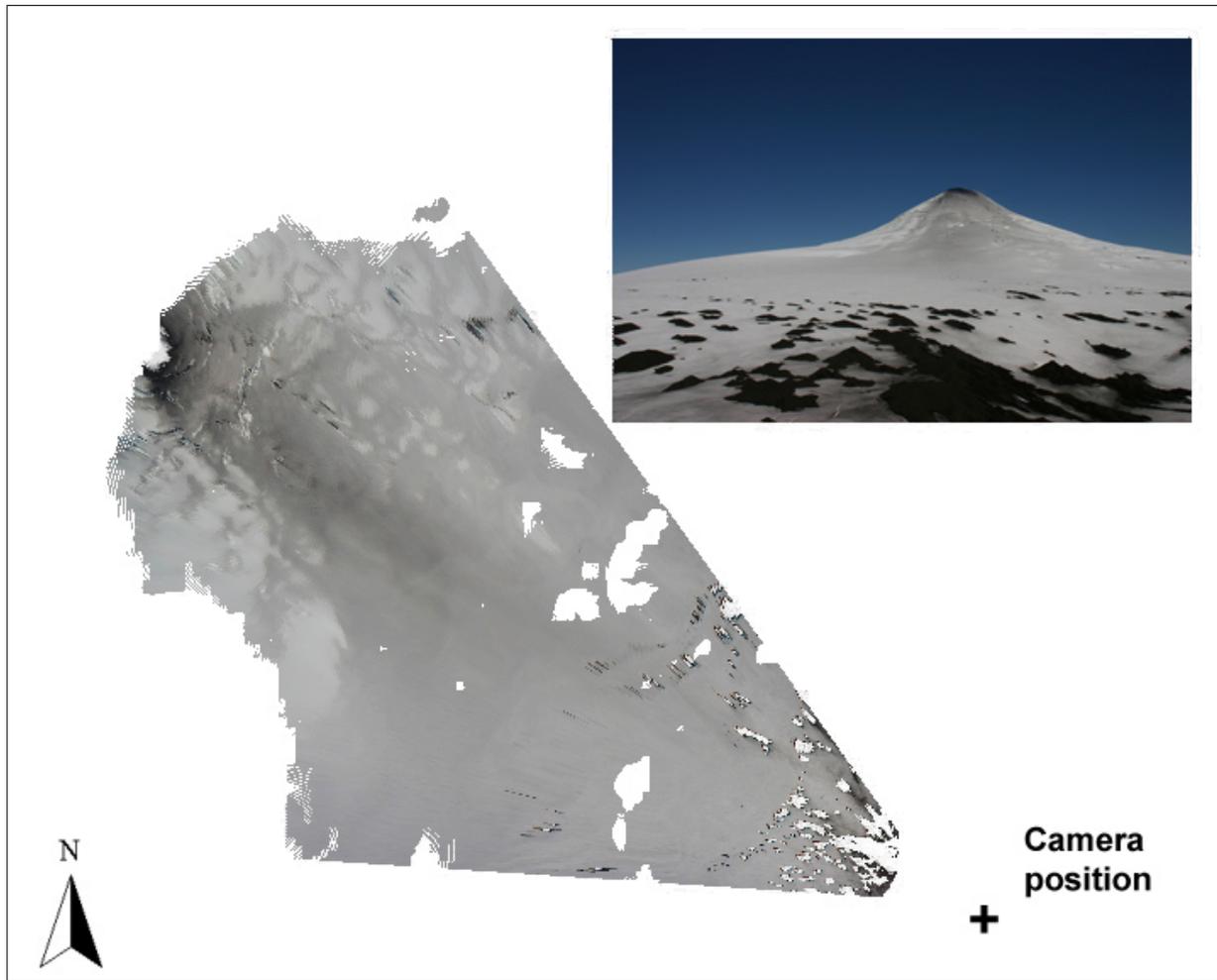
**Fig. 3.** Location of GPS measurements and the AWS in winter and summer. The lower panel shows the location of the flags installed on the glacier surface which were measured with GPS whilst a photograph was acquired for each site



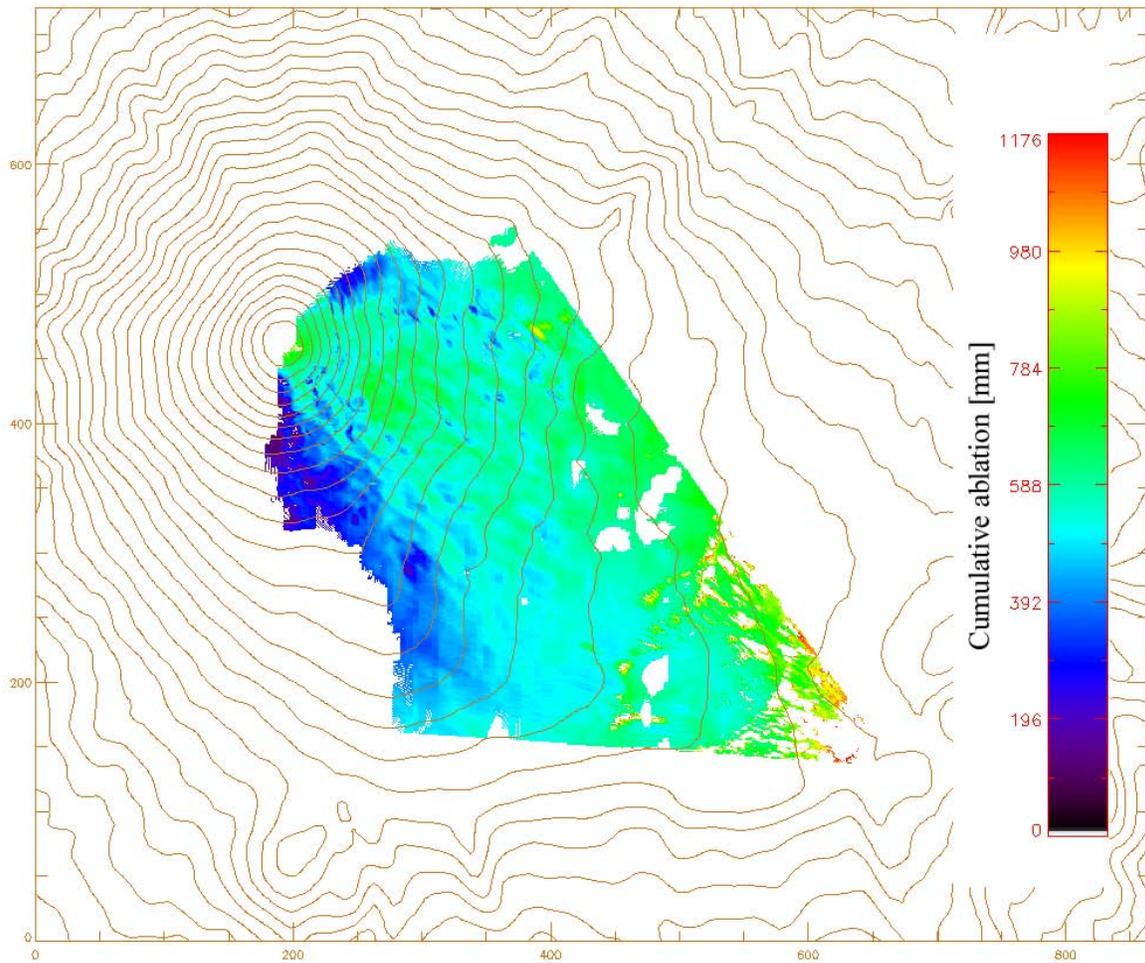
**Fig. 4.** Image of Volcán Villarrica and superimposed perspective projection of the digital elevation model. Every red dot corresponds to a grid cell in the DEM. Circles in the insets show the location of precise ground control points, which in the terrain were marked by flags visible in the original high resolution photograph



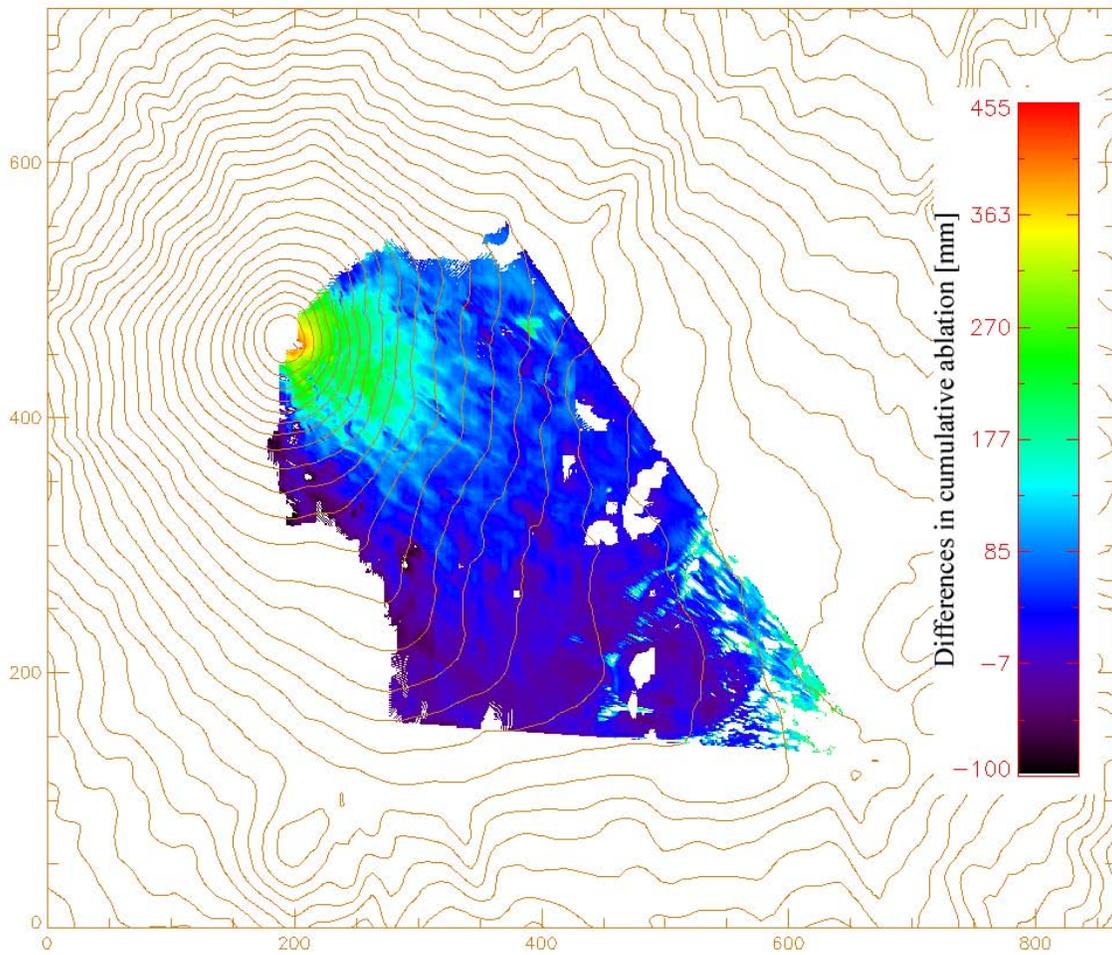
**Fig. 5.** Detail of two different images showing the changing position of crevasses on the upper section of the volcano. Left image is 25 March, 2006, right image is 14 January, 2007. Black are areas non visible from the standpoint of the camera.



**Fig. 6.** Pattern of tephra dispersion on the Volcán Villarrica, 25/12/2005. The inset shows the original image. The conical shape of the georeferenced image is due to the field of view of the camera, located on the lower right corner. There is a thick tephra layer around the crater rim and a band of darker snow following the prevailing W and NW winds. White areas in the georeferenced image are not visible from the camera point of view.



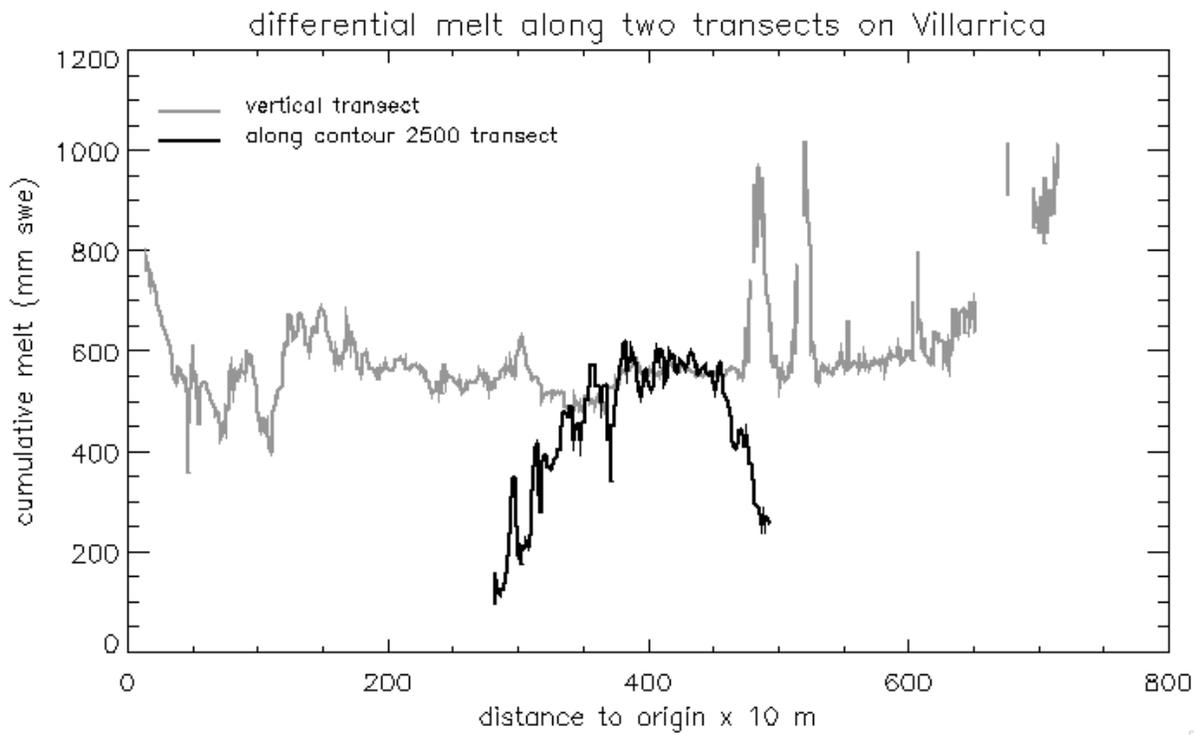
**Fig. 7.** Map of cumulative ablation from 5 to 16 January, 2007, modelled with albedo derived from daily photographs of the glacier (see Figure 6). The spatial variations are due to the pattern of tephra dispersion, which is more concentrated leeward of the prevailing W and NW winds. Axes units are 10 m.



**Fig. 8.** Map of ablation differences between the observed ash-covered volcano and a hypothetically clean snow surface with albedo parameterised according to Brock and others (2000).



**Fig. 9.** Photograph of Volcán Villarrica in spring, showing an irregular snow surface, which is likely to be a combination of ice flow and crevassing together with differential ablation.



**Fig. 10.** Profiles of cumulative ablation along the 2500 elevation contour (black) and across a vertical profile from the crater down in direction SE (grey). The origin of the black transect is the NE direction of the 2500 contour, opposite to the viewing direction of the observer. The origin of the vertical transect (grey) is the summit crater.

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